

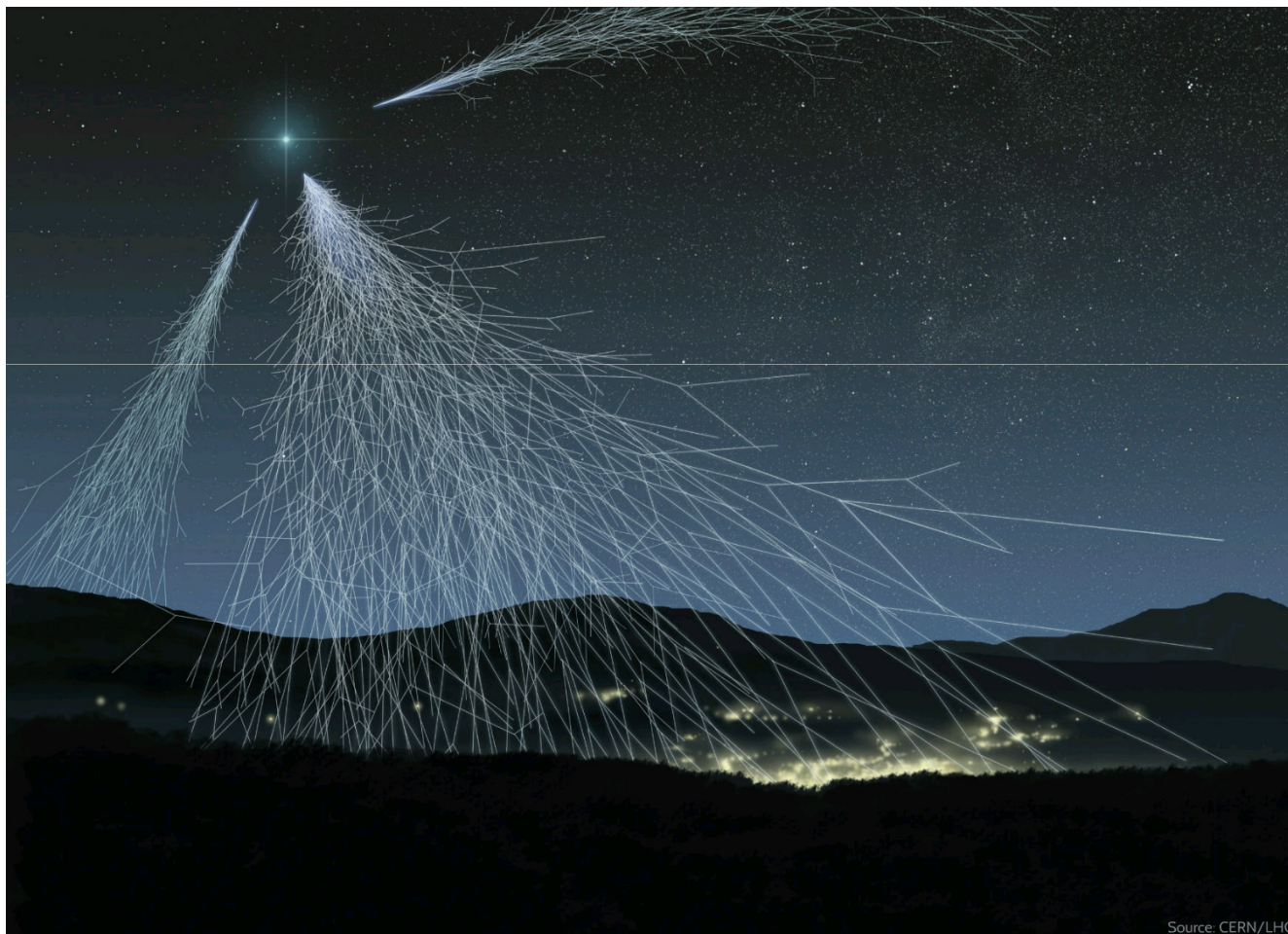
# Influence of total biomass and rainfall interception on cosmic-ray soil moisture measurements

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# Cosmic-ray soil moisture sensing

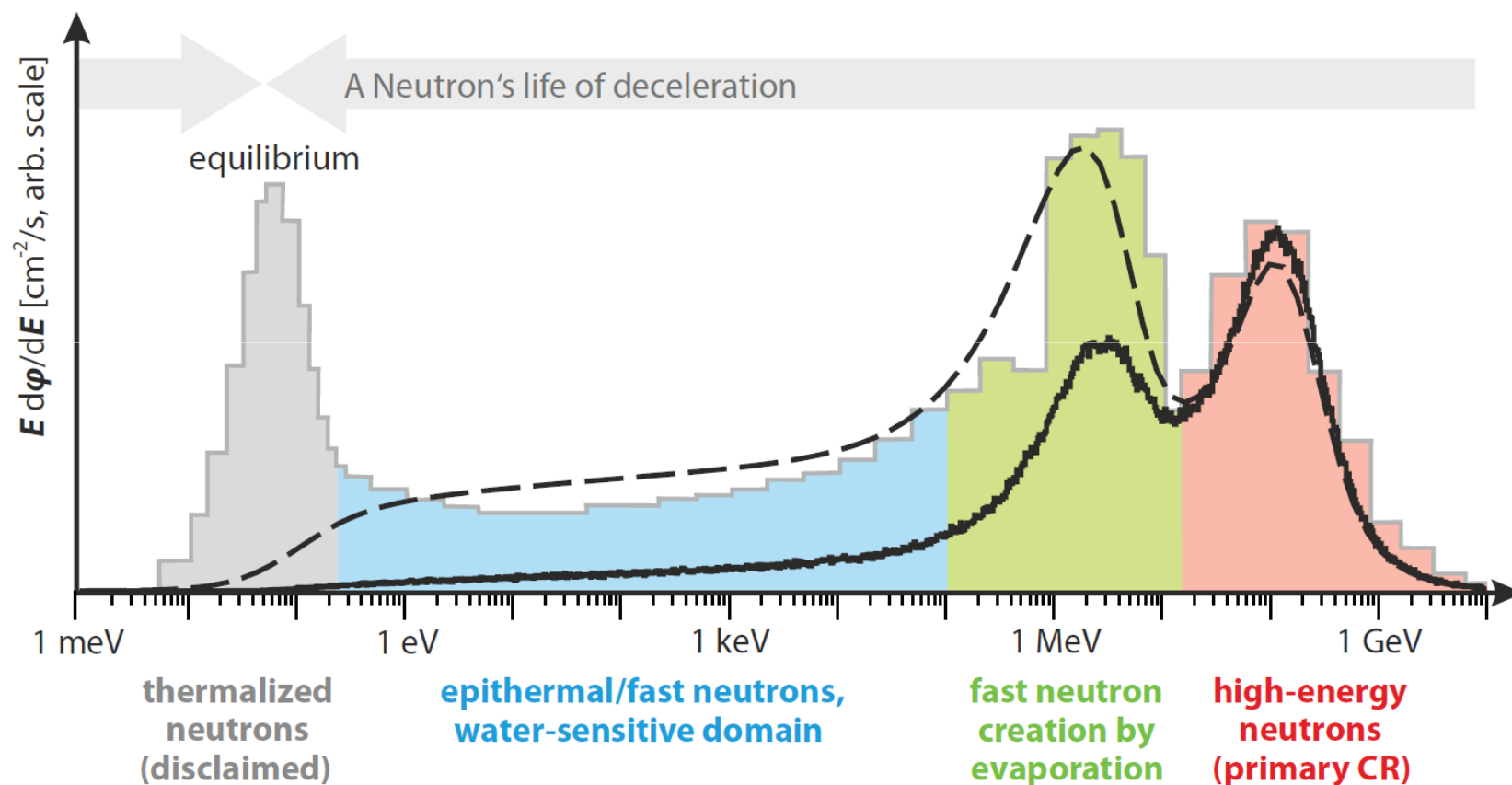
- Cosmic rays lead to emission of fast neutrons into the lower atmosphere



Source: CERN/LHC

# Cosmic-ray soil moisture sensing

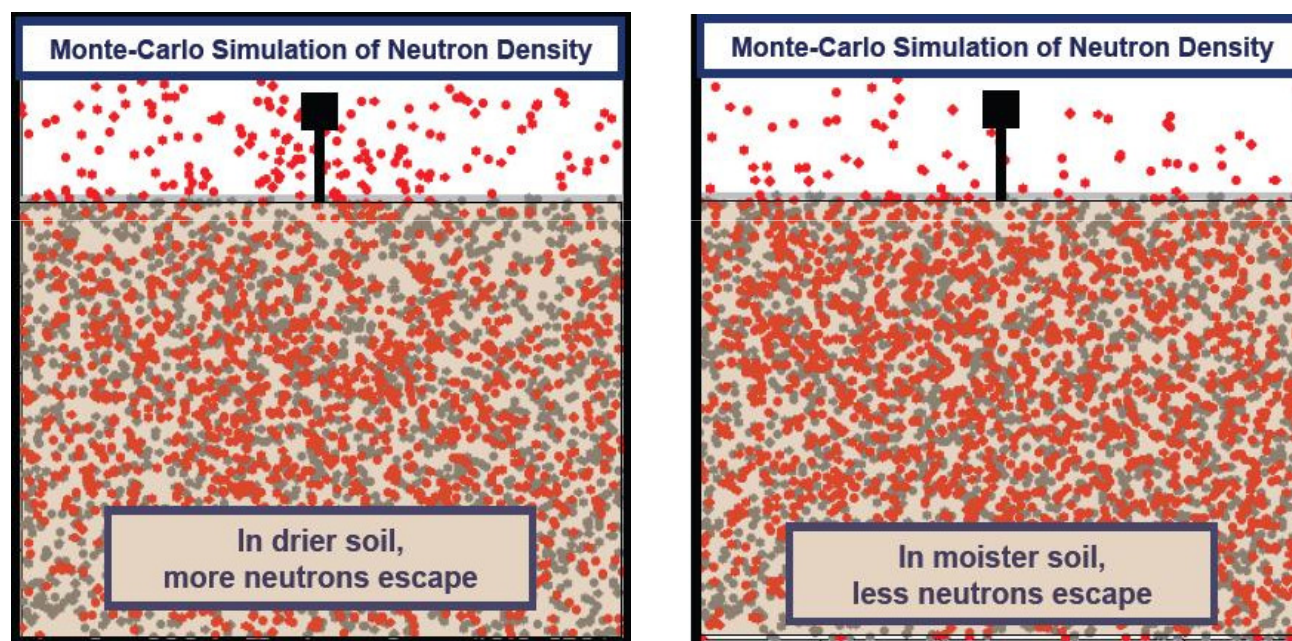
- Cosmic rays lead to emission of fast neutrons into the lower atmosphere



Neutron energy spectra at the surface (Köhli et al., 2015)

# Cosmic-ray soil moisture sensing

- Cosmic rays lead to emission of fast neutrons into the lower atmosphere
- Fast neutrons lose energy primarily through collisions with hydrogen mainly located in the soil





# Cosmic-ray soil moisture sensing

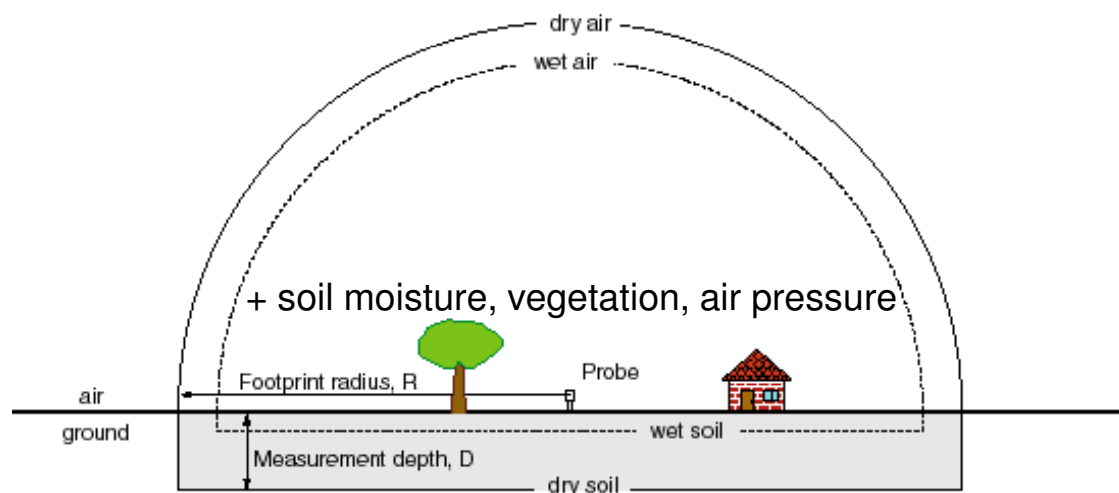
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- Continuous measurement of neutron flux to estimate soil moisture at hourly resolution



Cosmic ray probe in the field of Rollesbroich (May 2011)

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- Large Footprint (~150-250 m radius, ~5-70 cm depth)



Zreda et al., 2012, HESS  
Köhli et al., 2015, WRR



Cosmic ray probe in the field  
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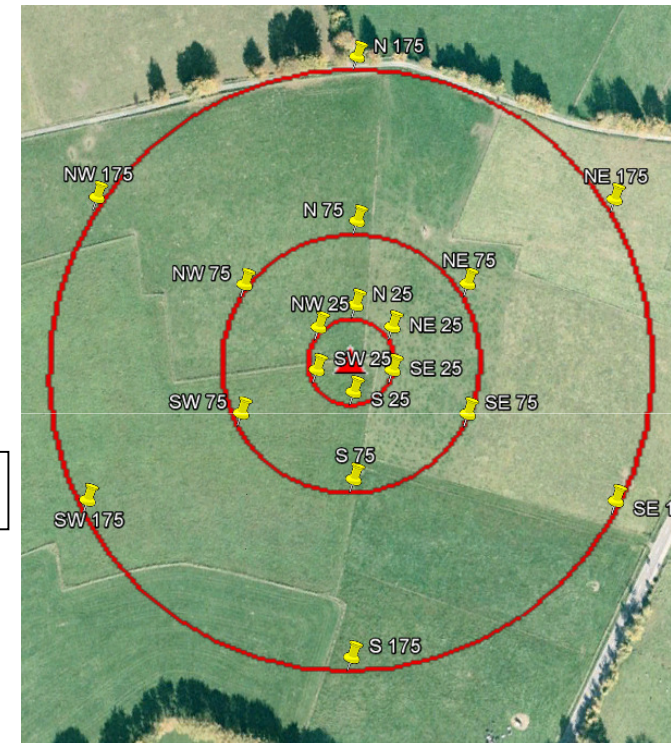
# Cosmic-ray soil moisture sensing

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- Fast neutrons lose energy primarily through collisions with hydrogen mainly located in the soil
- Continuous measurement of neutron flux to estimate soil moisture at hourly resolution
- Large Footprint (~150-250 m radius, ~5-70 cm depth)
- Simple calibration with the N0-method:

$$\theta_{tot} = \frac{\rho_{bd} \cdot a_0}{(N_{corr} / N_0) - a_1} - \rho_{bd} \cdot a_2$$

$$\theta_{vol} = \theta_{tot} - \theta_{lat} - \theta_{SOM}$$

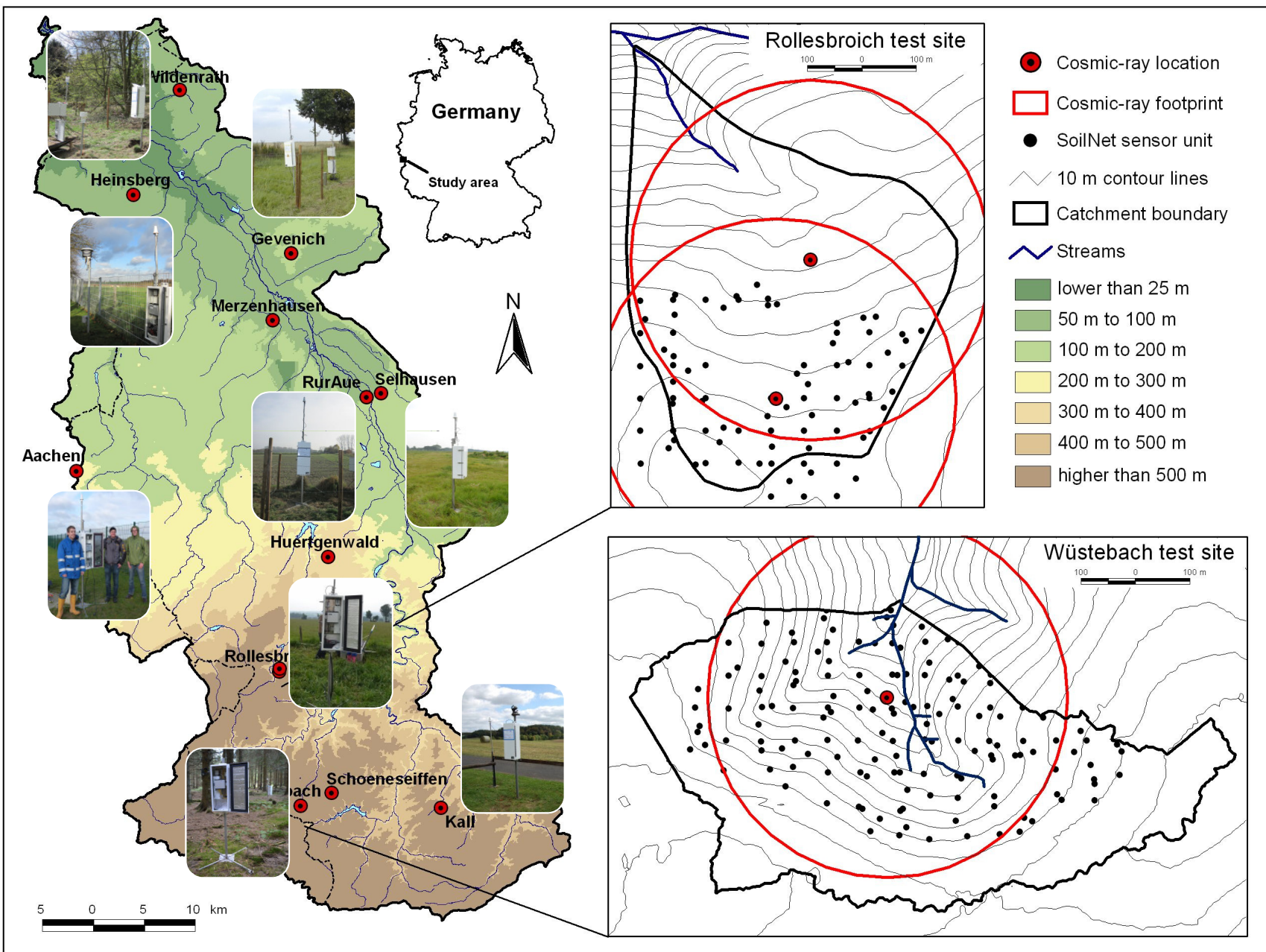
$\theta_{tot}$	Total water content [cm <sup>3</sup> /cm <sup>3</sup> ]
$a_0, a_1, a_2$	Constant parameters
$\rho_{bd}$	Soil density [g/cm <sup>3</sup> ]
$N_{corr}$	Measured and corrected fast neutrons (counts/h)
$N_0$	Fast neutron counts for dry soil conditions (counts/h)
$\theta_{vol}$	Volumetric soil water content [cm <sup>3</sup> /cm <sup>3</sup> ]
$\theta_{lat}$	Lattice water content [cm <sup>3</sup> /cm <sup>3</sup> ]
$\theta_{SOM}$	Water equivalent of soil organic matter [cm <sup>3</sup> /cm <sup>3</sup> ]



Field calibration in Rollesbroich (May 2011)



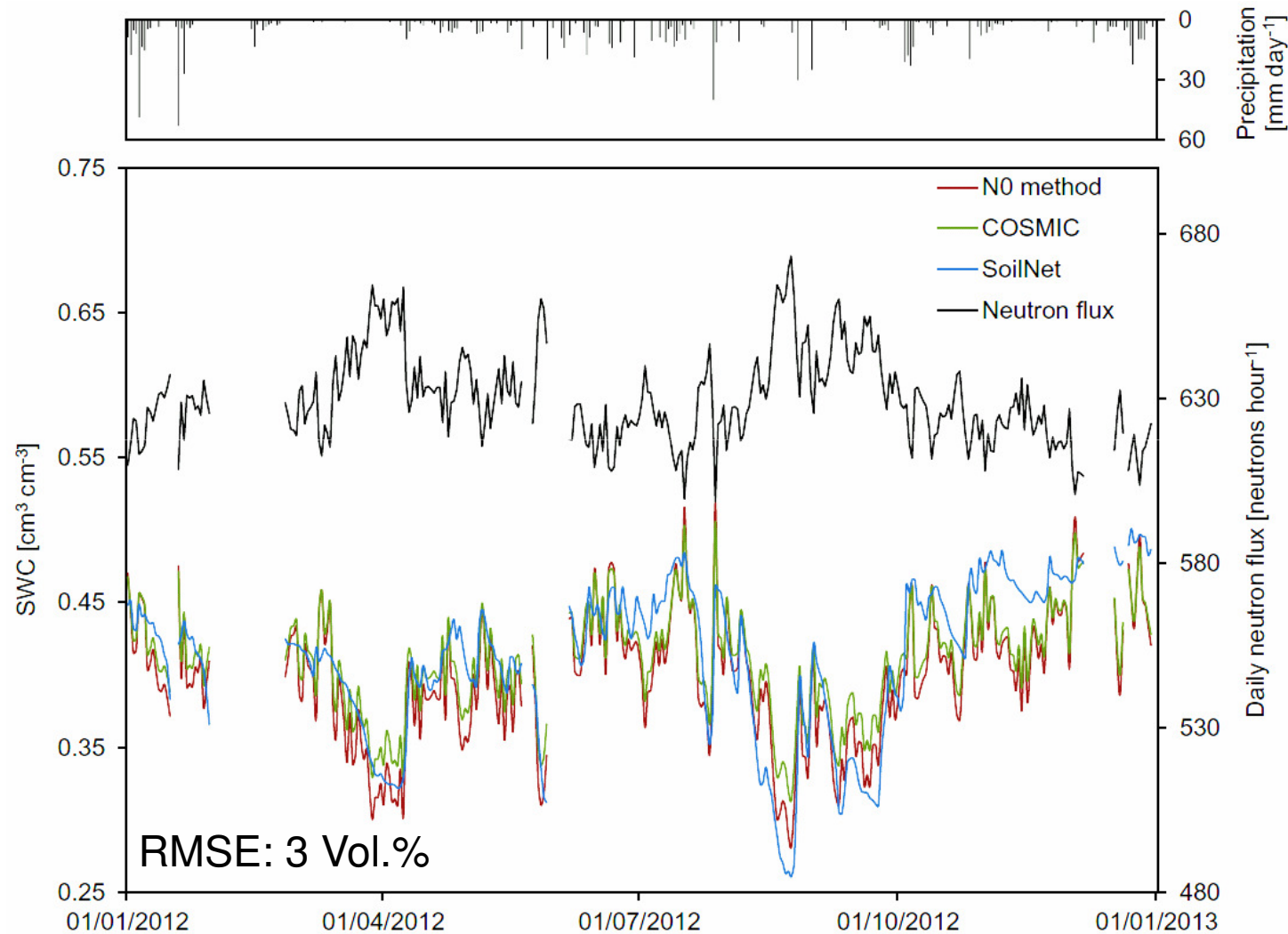
# Network of 13 cosmic-ray soil moisture probes





# Comparison with in-situ soil moisture data

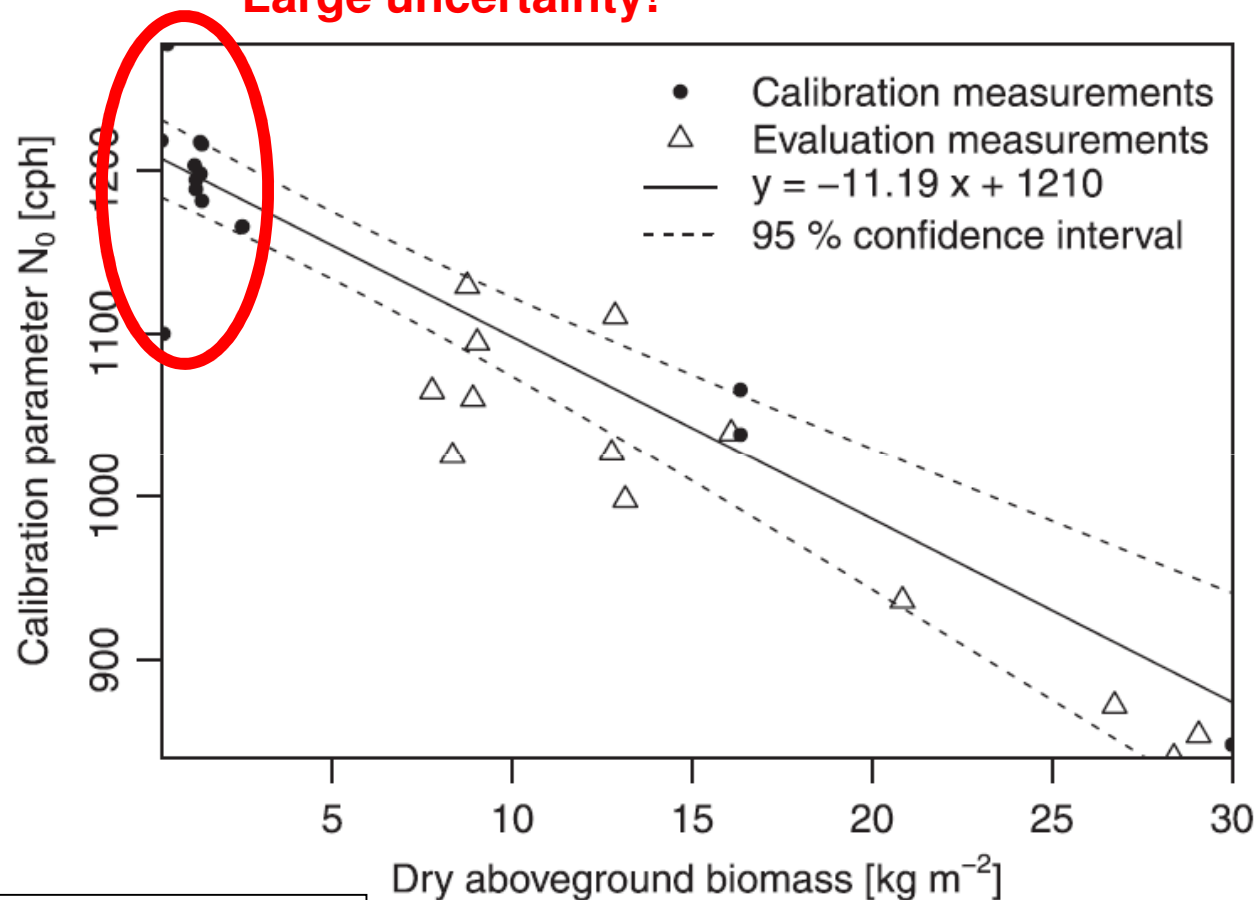
Example: grassland test site Rollesbroich



Baatz et al. (2014), J. Hydrol.

# Cosmic-ray data is influenced by biomass

Large uncertainty!



$$\theta_{tot} = \frac{\rho_{bd} \cdot a_0}{(N_{corr} / N_0) - a_1} - \rho_{bd} \cdot a_2$$

Baatz et al. (2015), WRR

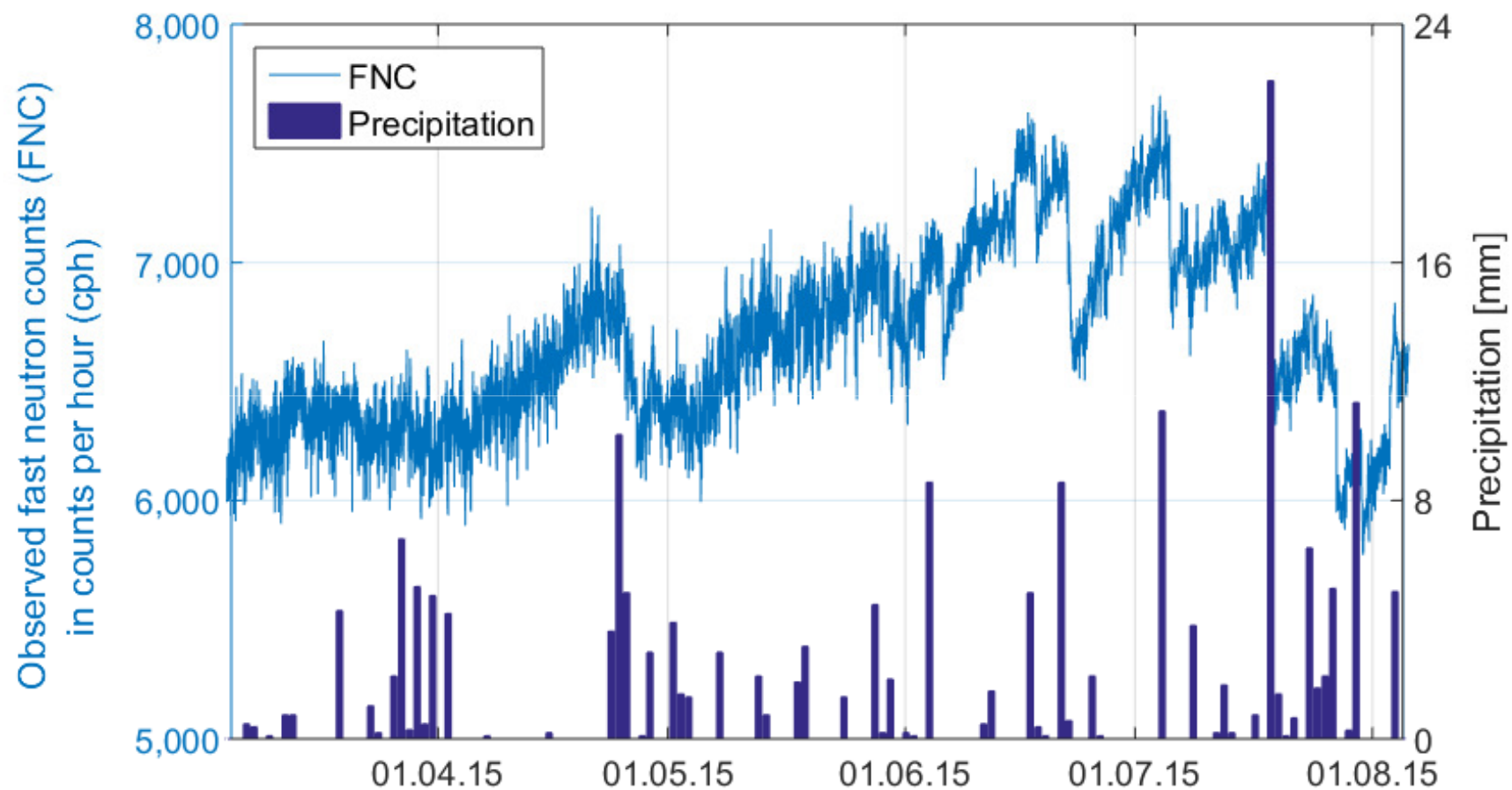
## Selhausen test site: Experimental setup



Seven cosmic-ray neutron probes (4 CRS1000, 2 CRS2000/B, 1 cross calibrator)

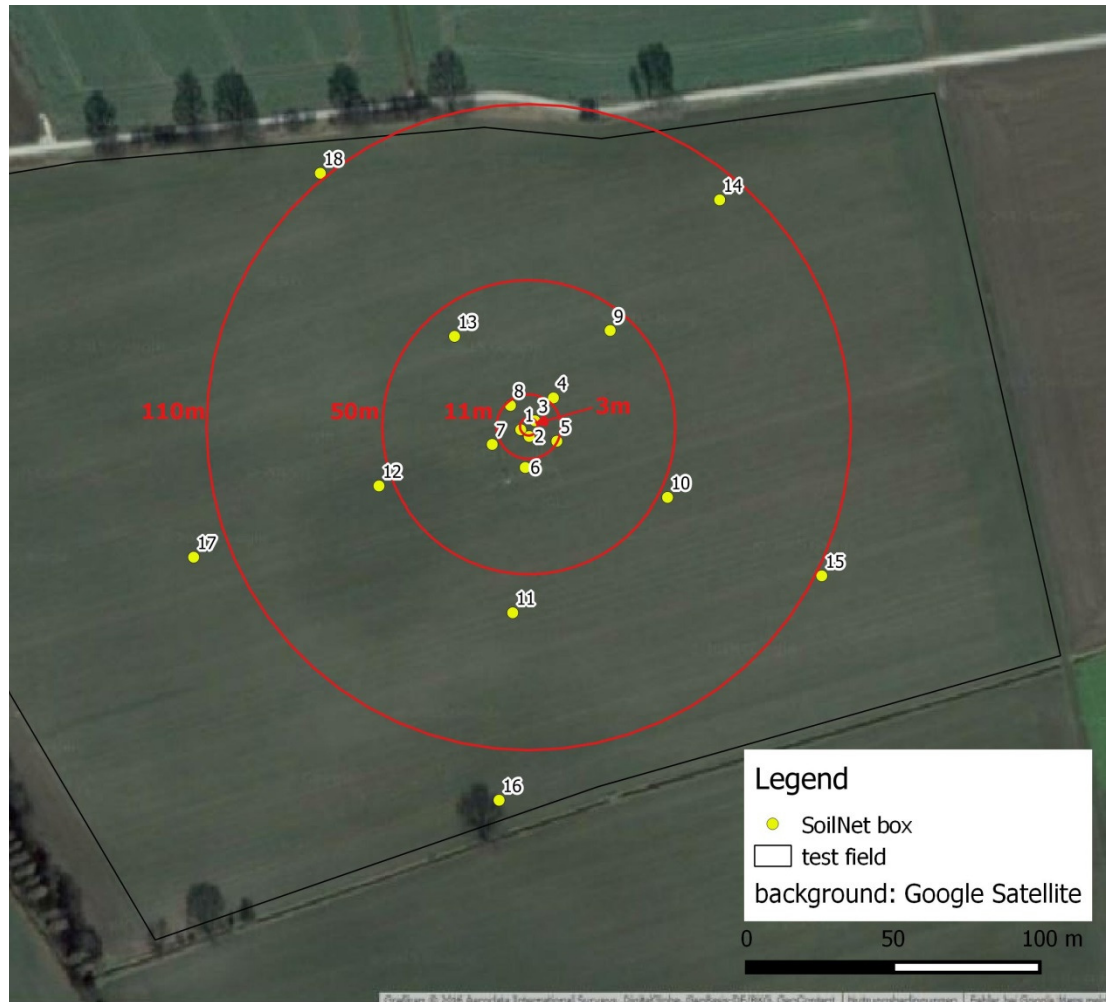


# Measured fast neutron intensities





# Experimental setup: SoilNet

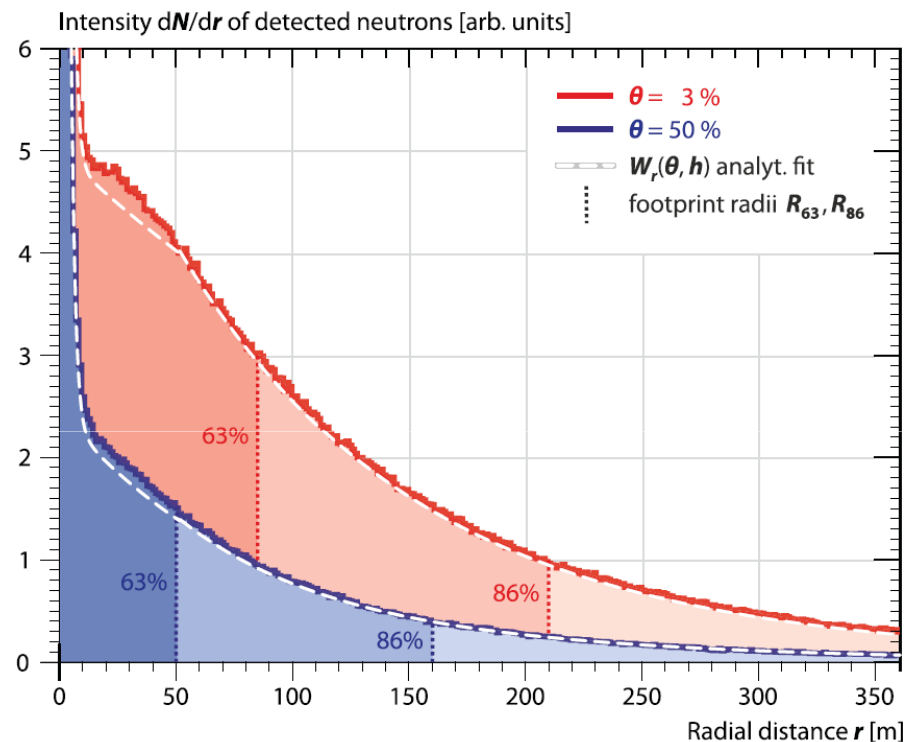


SoilNet sensor unit



18 SoilNet sensor units each with 2 SMT100 SWC sensors installed at 5, 10 and 20 cm depth

# Comparison of Cosmic-ray with SoilNet data



Köhli et al., 2015

→ SWC measured by SoilNet has to be weighted accordingly to be comparable to CRNP

*Vertically → depending on penetration depth, and soil water content*

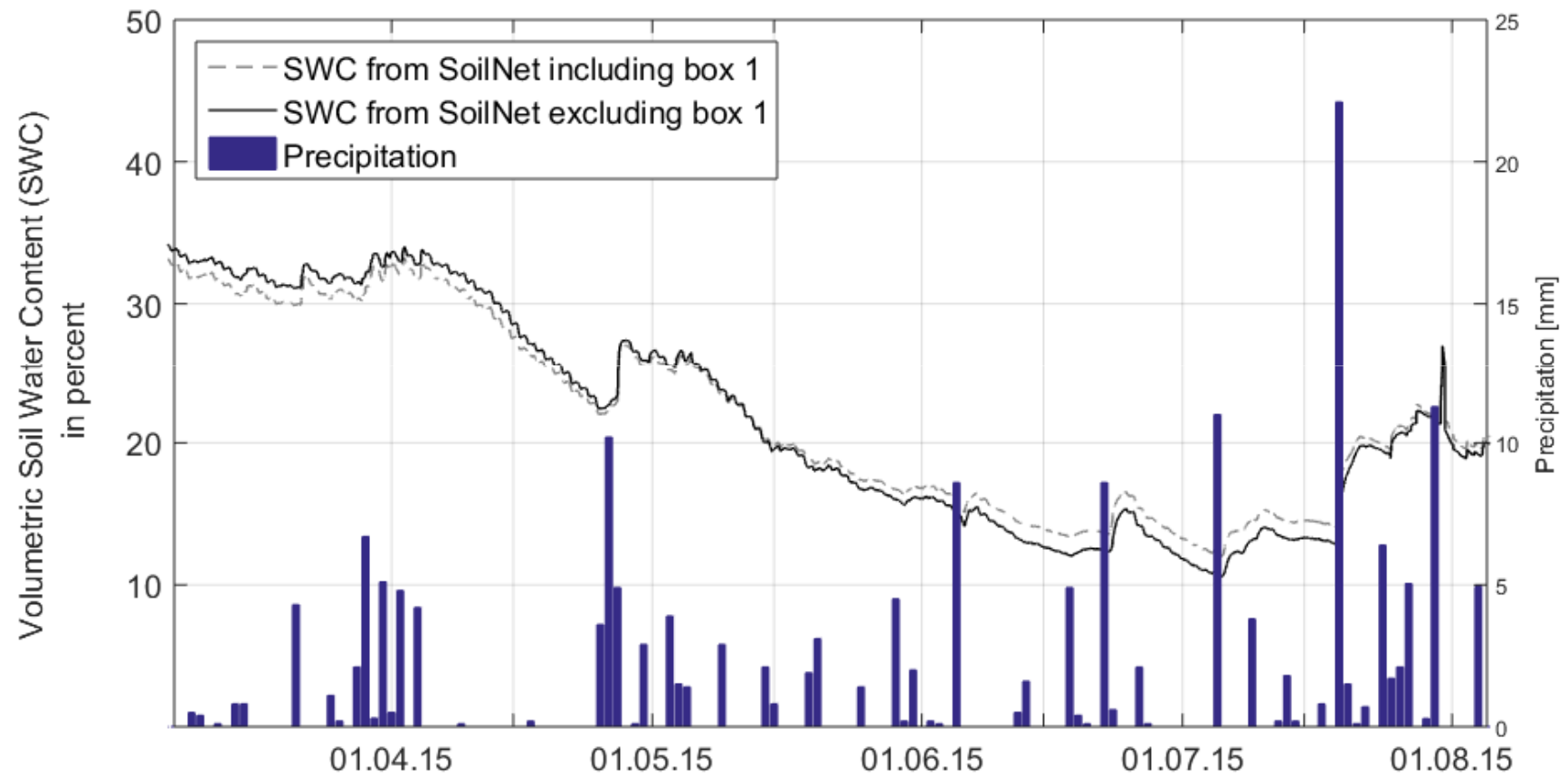
$$D_{86}(r, \theta) = \rho_{bd}^{-1} \left( p_0 + p_1 \left( p_2 + e^{-r/100} \right) \frac{p_3 + \theta}{p_4 + \theta} \right)$$

$$W_d(r, \theta) \propto e^{-2d/D_{86}(r, \theta)}$$

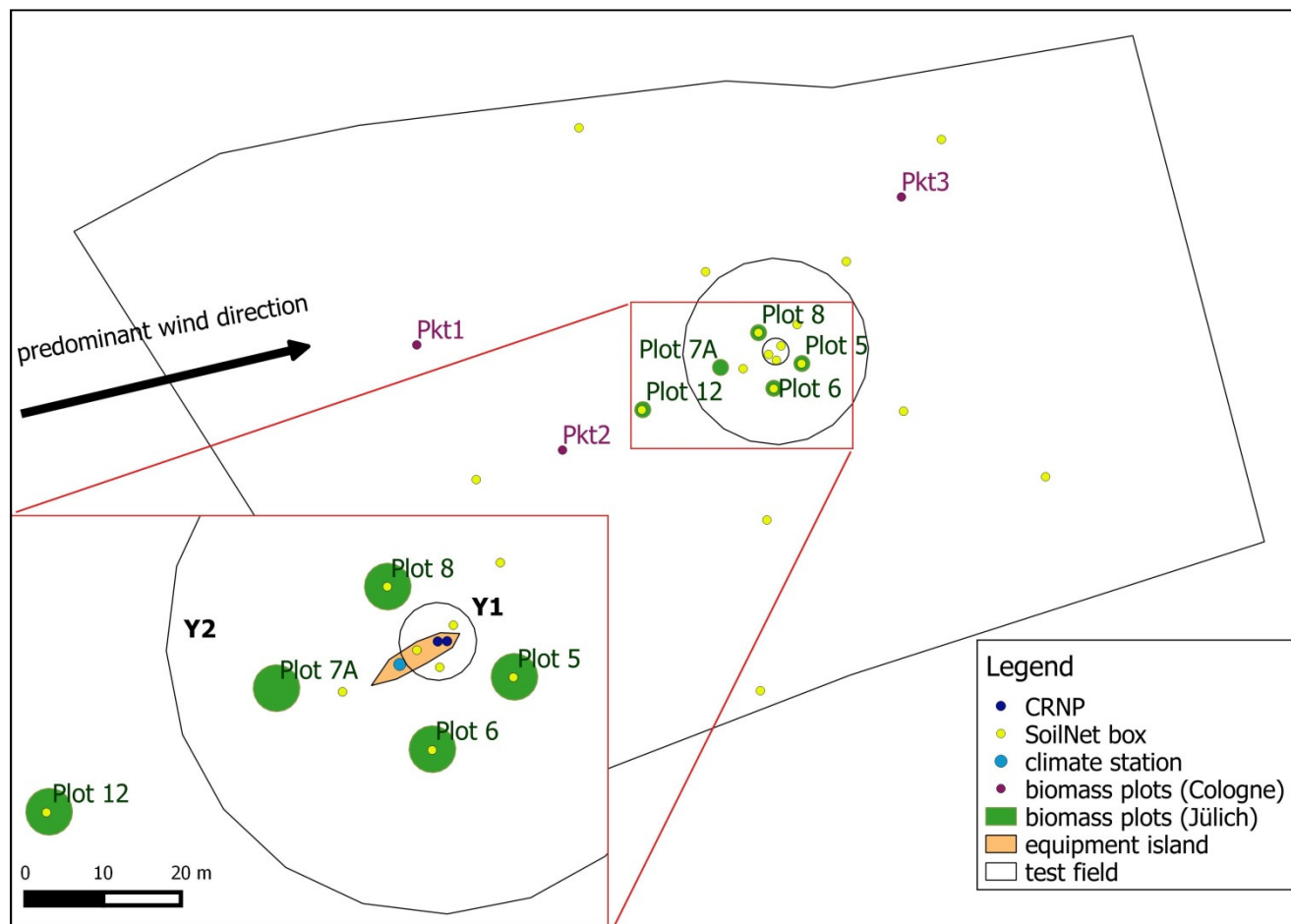
*Horizontally → depending on distance, soil water content and air humidity*

$$W_r(h, \theta) \approx \begin{cases} F_1 e^{-F_2 r} + F_3 e^{-F_4 r}, & 0.5 \text{ m} < r \leq 50 \text{ m} \\ F_5 e^{-F_6 r} + F_7 e^{-F_8 r}, & 50 \text{ m} < r < 600 \text{ m} \end{cases}$$

# Weighted soil moisture



# Biomass, LAI and interception sampling according to ICOS protocols



Biomass sampling plots



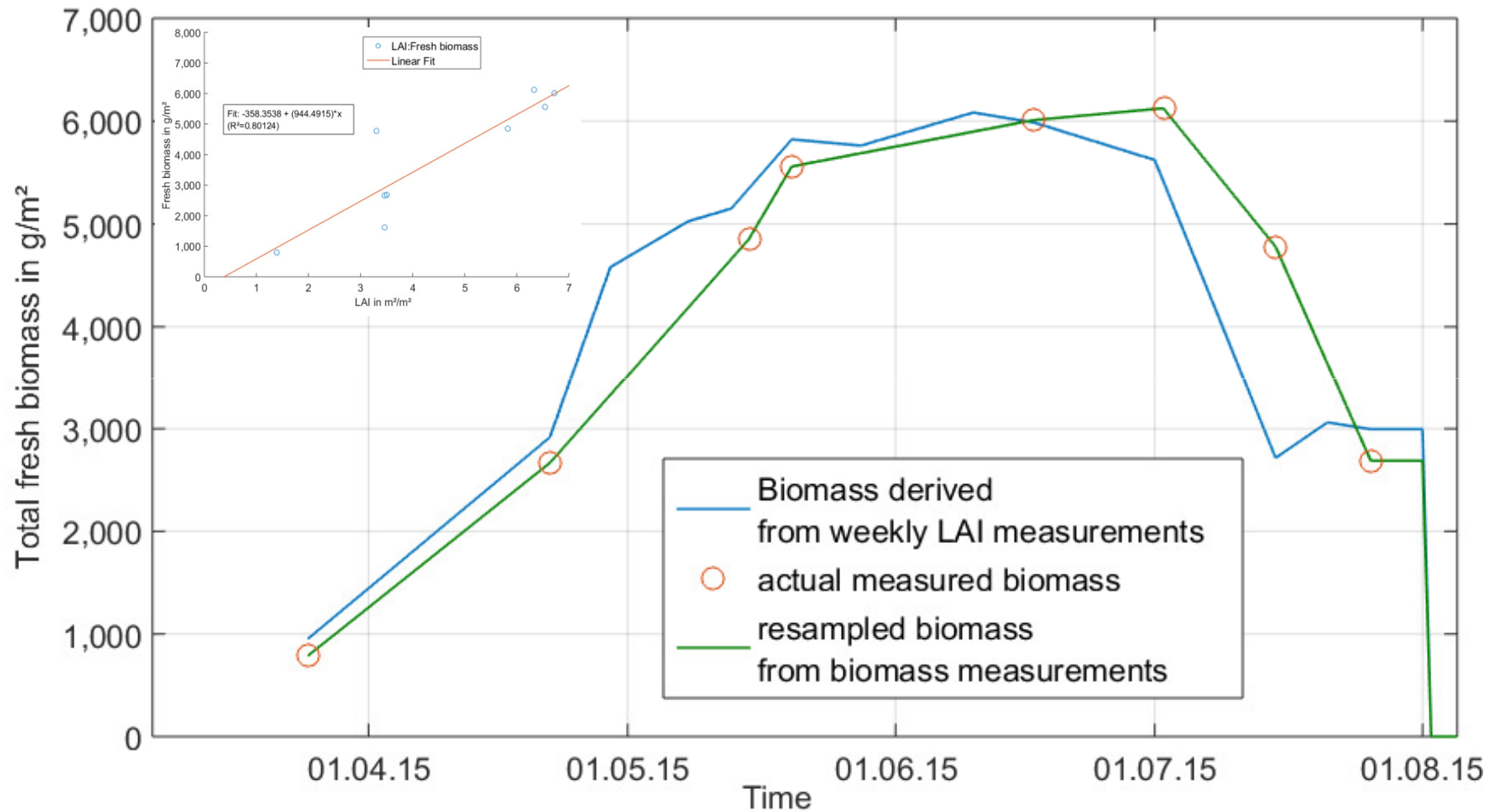
LAI measurements



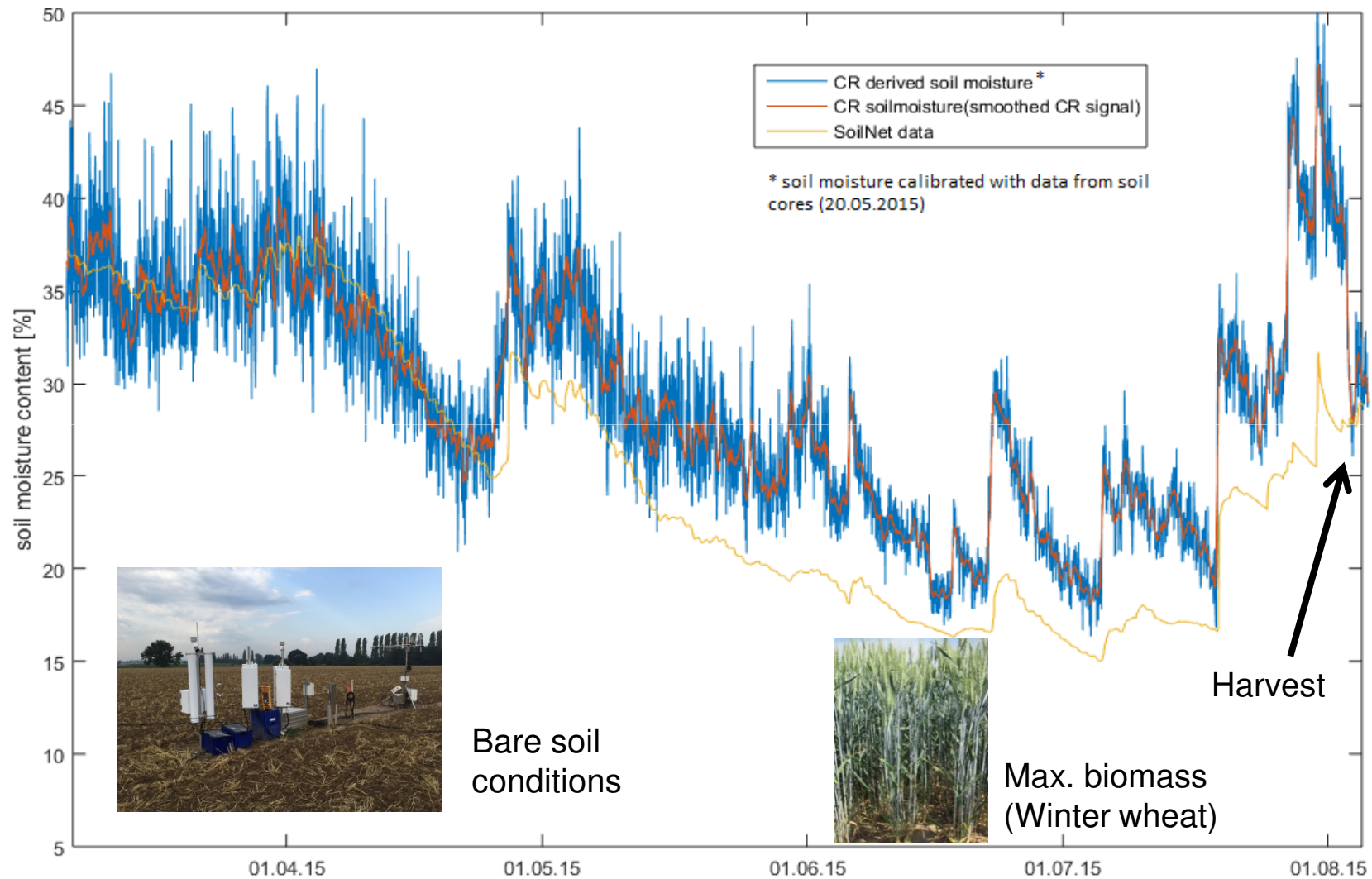
Interception measurements



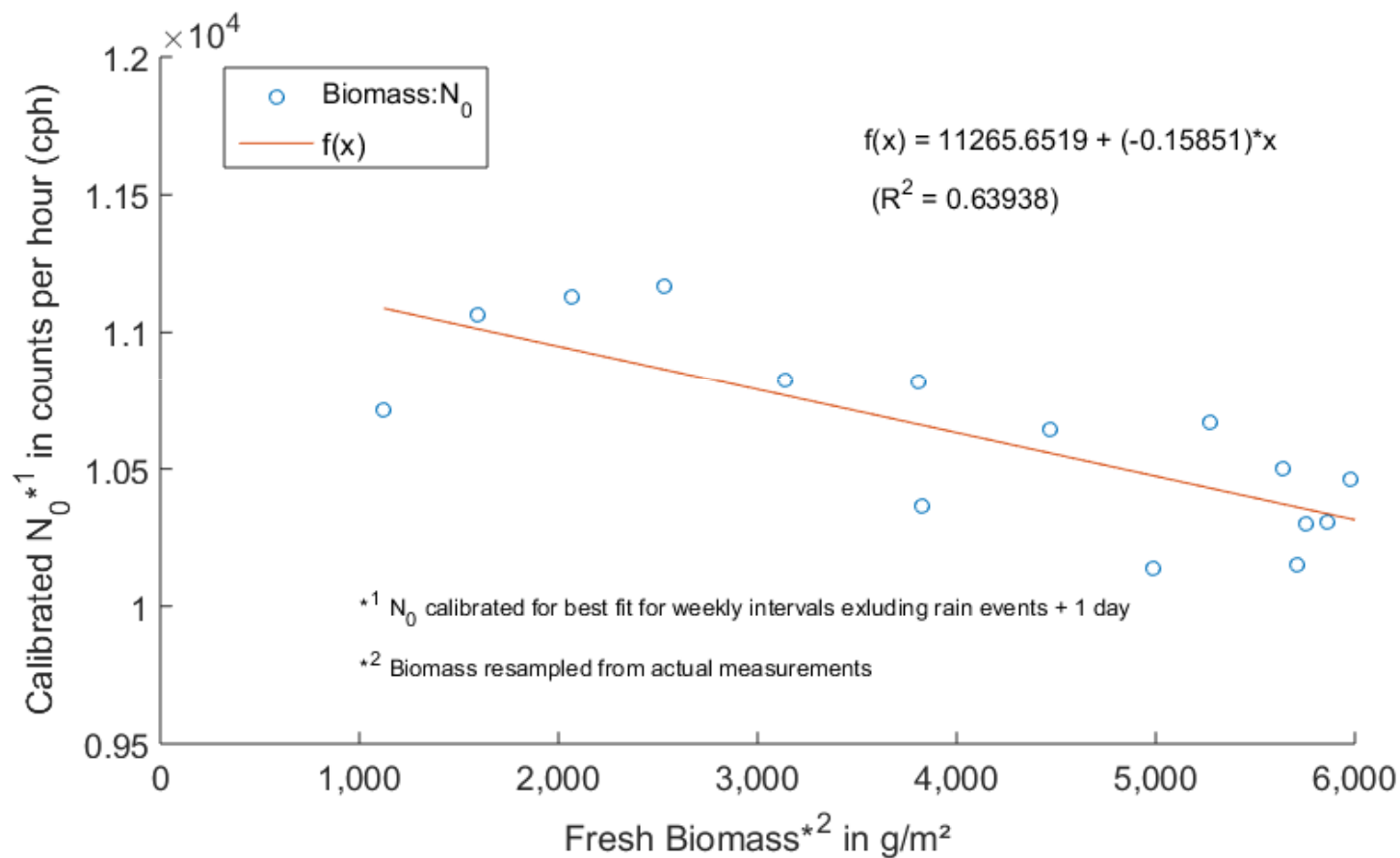
# Biomass development



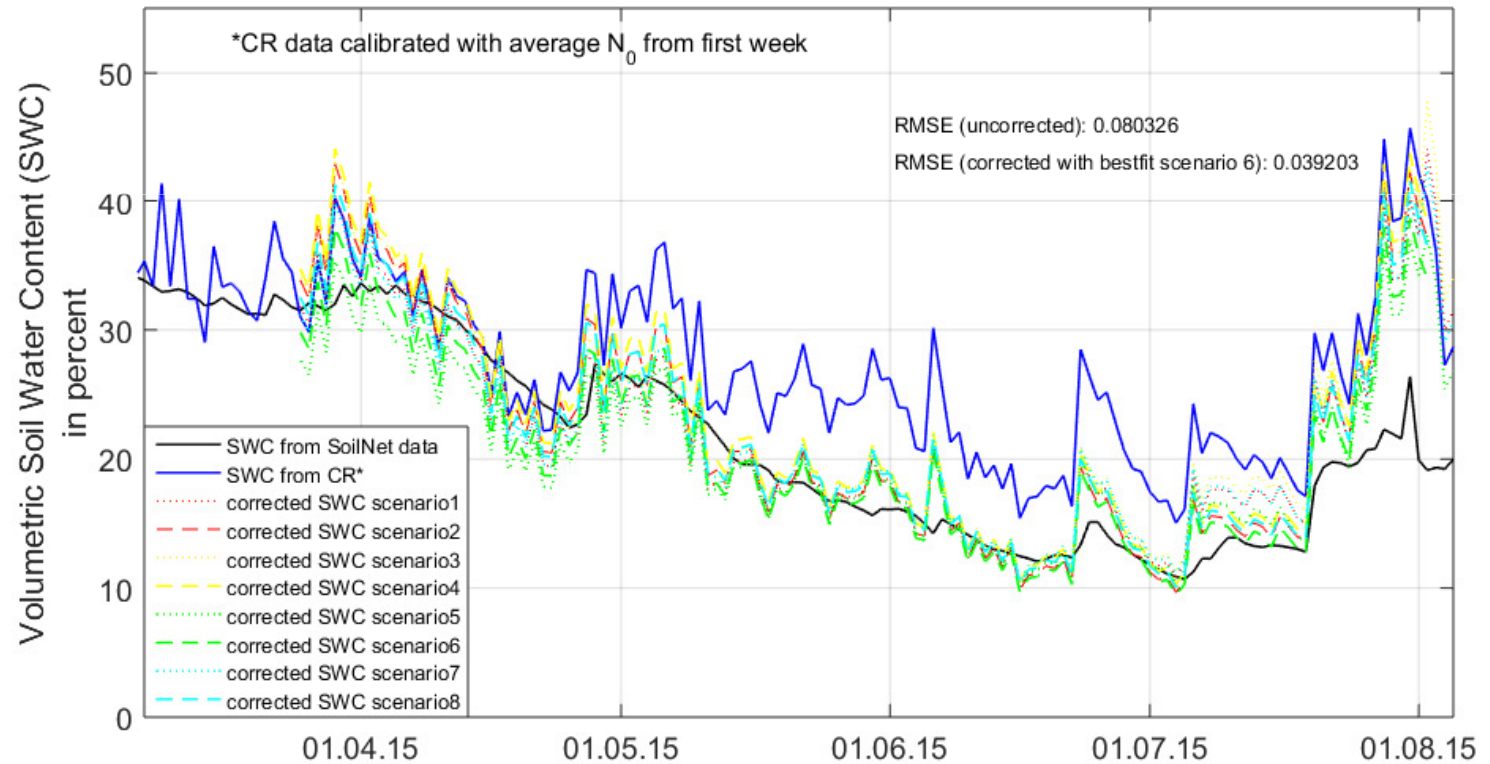
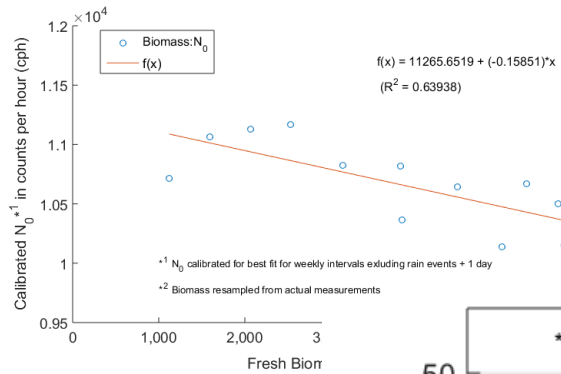
# Biomass effects on cosmic-ray soil moisture



# Relationship between N0 and biomass

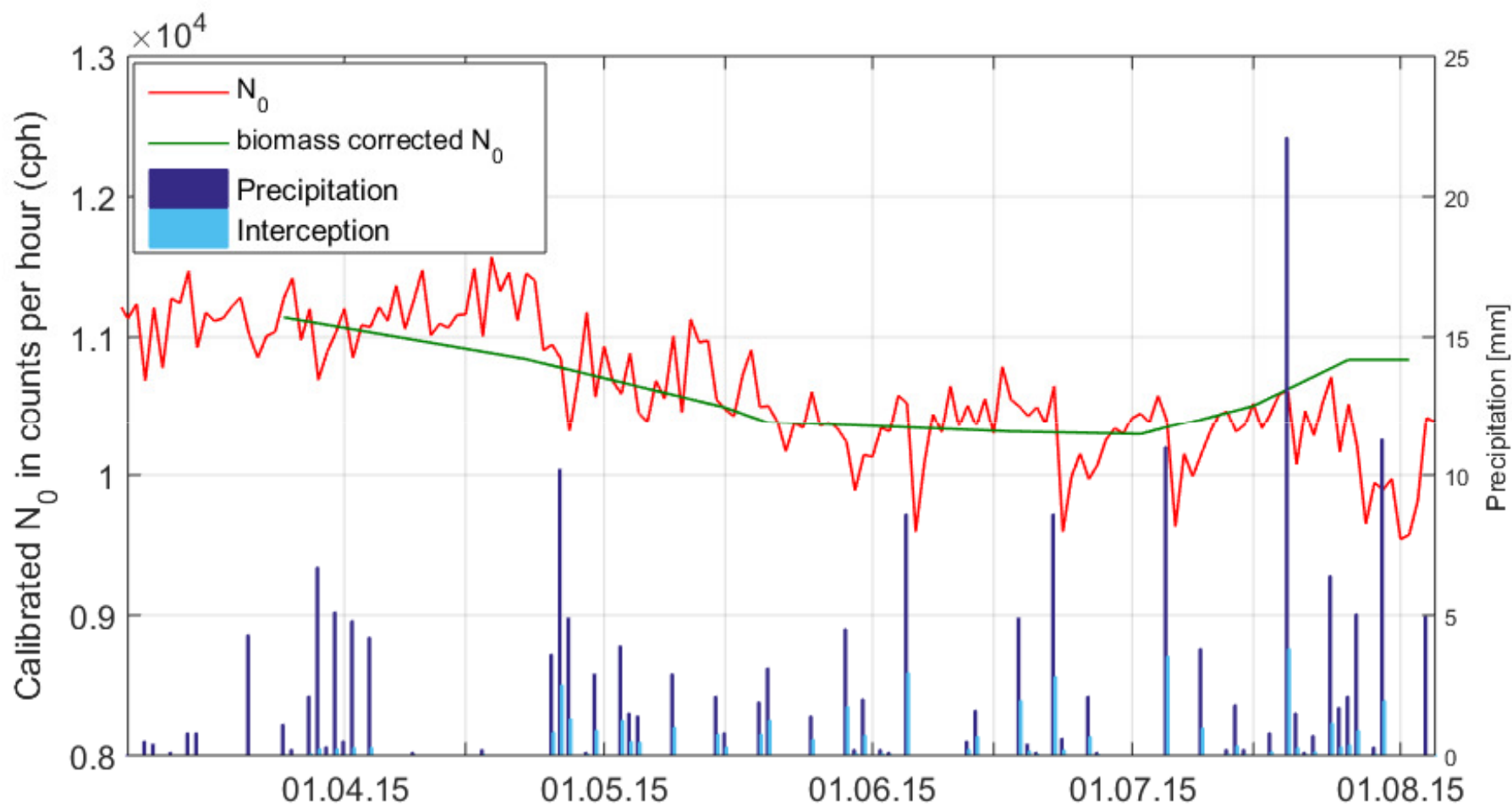


# Correction of biomass effects

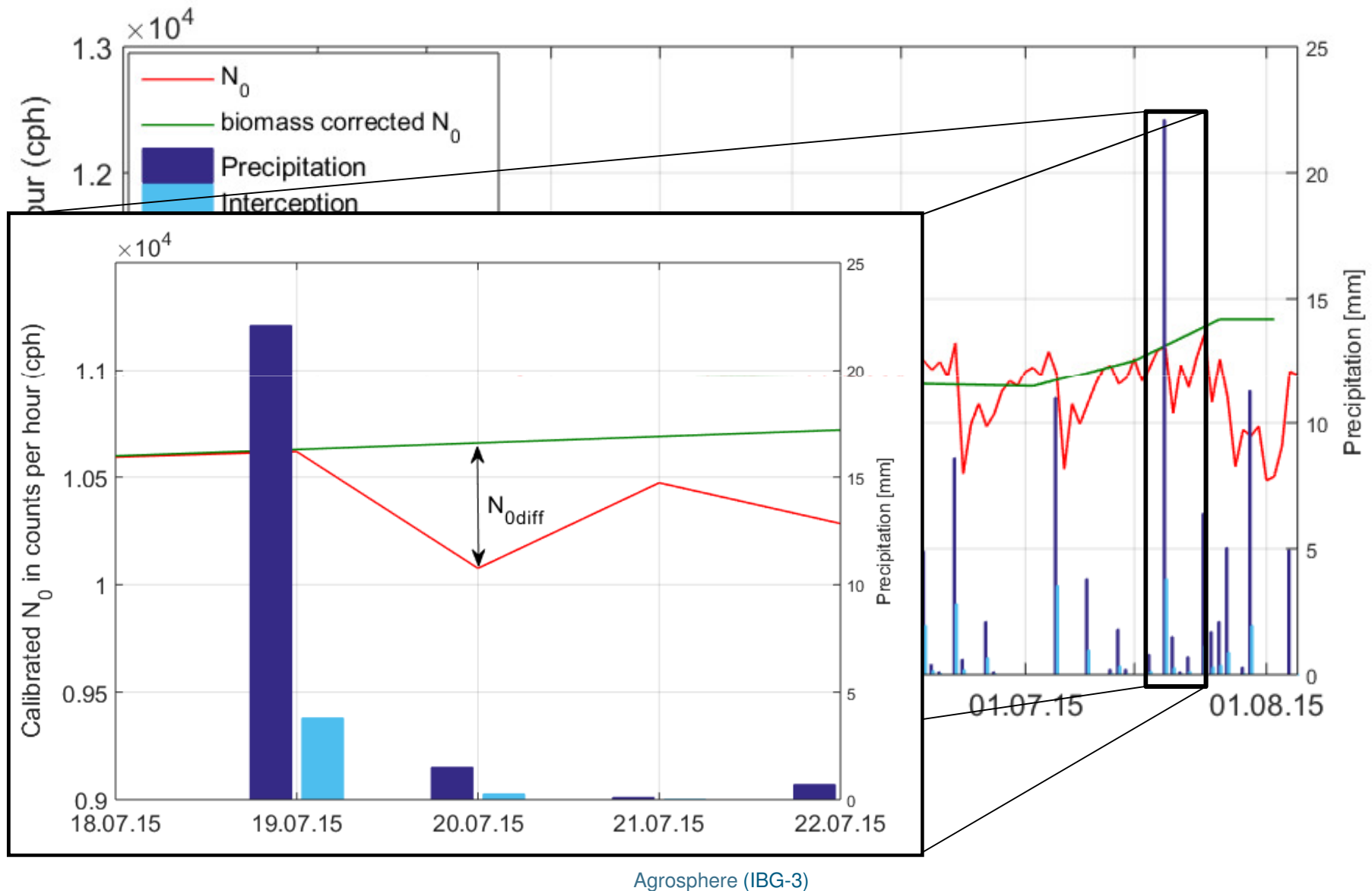




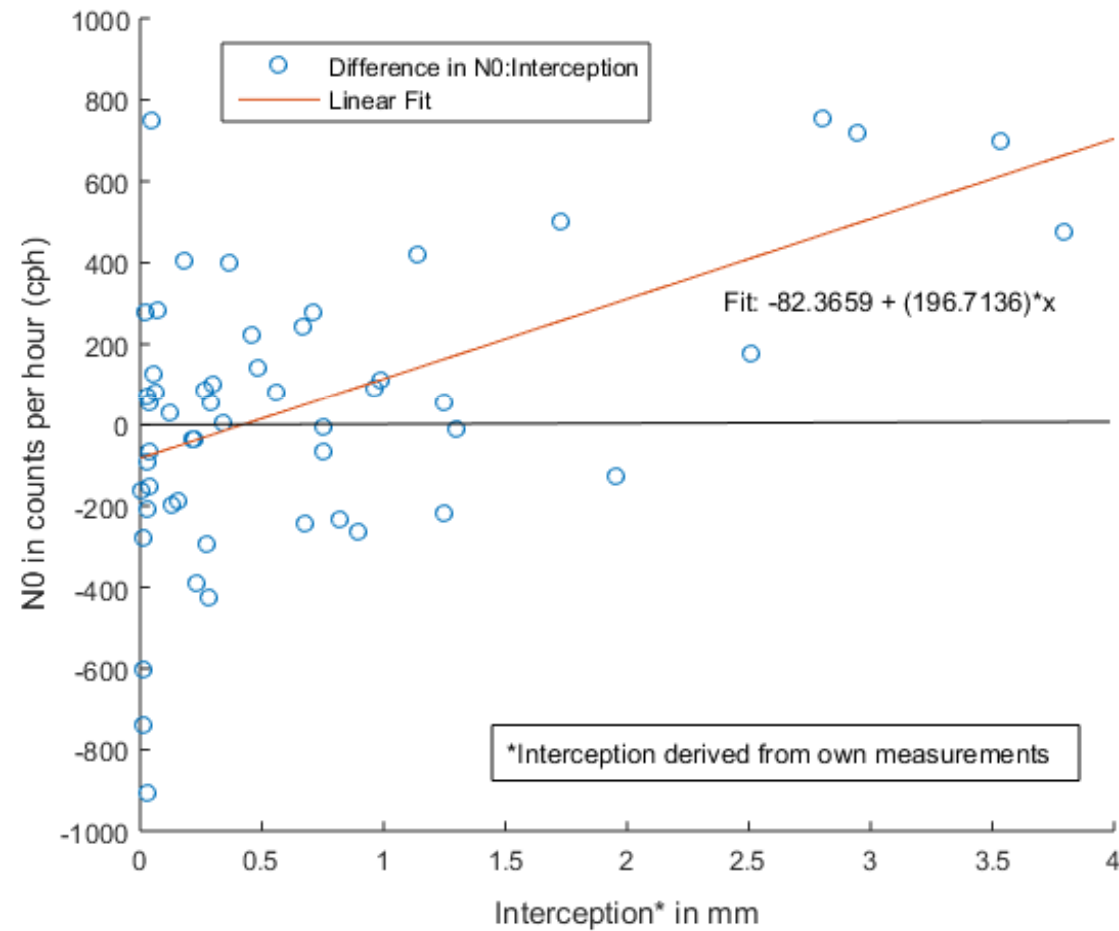
# Influence of interception effects on N<sub>0</sub>



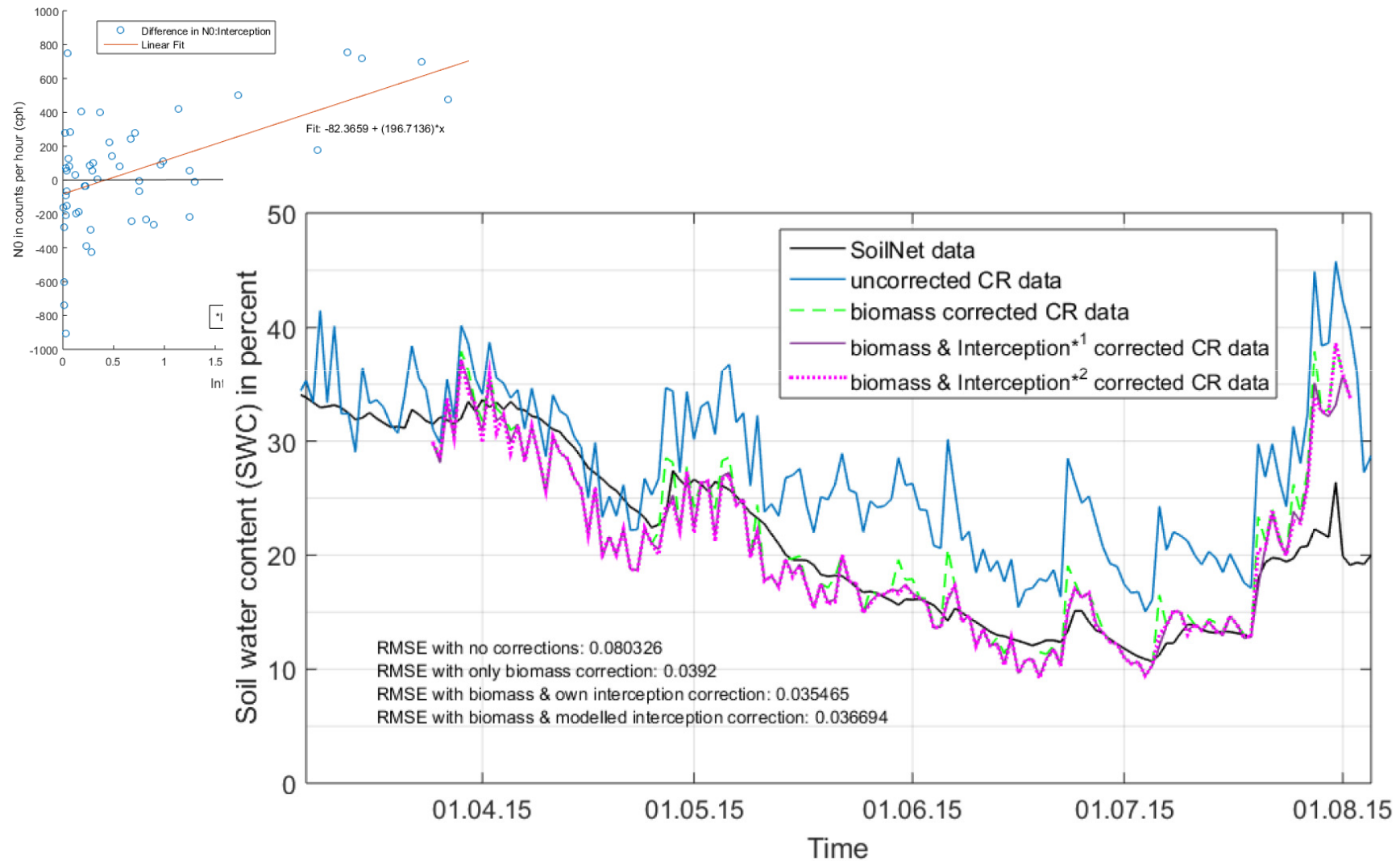
# Influence of interception effects on $N_0$



# Influence of interception effects on N0



# Influence of interception effects on N0





## Conclusions

- This study stressed the importance of biomass and interception effects on cosmic-ray soil moisture measurements in agricultural environments
- The RMSE could be reduced from  $0.08 \text{ m}^3/\text{m}^3$  to  $0.035 \text{ m}^3/\text{m}^3$  by applying corrections for biomass and interception effects
- Remaining deviations are likely due for strong soil moisture gradients in the top 5 cm of the soil
- Further experiments with different crops have started this year (e.g. sugar beet)

*Thanks a lot for your attention!*

